NUCLEAR CHIRAL THERMODYNAMICS and the

FUNCTIONAL RENORMALIZATION GROUP



Wolfram Weise

ECT* Trento and Technische Universität München

- Introductory glance at the **QCD phase diagram**
- Chiral models of the nuclear equation of state
- Beyond mean field: non-perturbative fluctuations and FRG approach
- Symmetric and asymmetric **nuclear matter**
- **Neutron matter** and **neutron stars**
- Density & temperature dependence of **chiral order parameter**

(based on doctoral thesis work of **Matthias Drews**)

NUCLEAR MATTER and QCD PHASES



- NN distance:
- energy per nucleon:
- compression modulus:

 $egin{aligned} k_F &\simeq 1.4~fm^{-1} \sim 2m_\pi \ d_{NN} &\simeq 1.8~fm \simeq 1.3~m_\pi^{-1} \ E/A &\simeq -16~MeV \ K &= (260 \pm 30)~MeV &\sim 2m_\pi \end{aligned}$



Nuclear Forces



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Empirical freeze-out systematics: thermal hadron production







• Chiral EFT represents QCD at energy/momentum scales ${f Q} << 4\pi\,{f f}_\pi \sim\,1\,{
m GeV}$

Strategies at the interface between QCD and nuclear physics :

In-medium **Ch**iral **P**erturbation **T**heory based on **non-linear sigma model** (with inclusion of nucleons)

expansion of free energy density in powers of Fermi momentum Chiral Nucleon-Meson model based on linear sigma model

non-perturbative Renormalization Group approach



PART I:

In-medium Chiral Perturbation Theory and the nuclear many-body problem





CHIRAL EFFECTIVE FIELD THEORY



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NUCLEAR INTERACTIONS from CHIRAL EFFECTIVE FIELD THEORY

Weinberg

Bedaque & van Kolck

Bernard, Epelbaum, Kaiser, Meißner; ...



Systematically organized HIERARCHY

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IN-MEDIUM CHIRAL PERTURBATION THEORY



NUCLEAR MATTER

 ${
m E_0/A} = -16\,{
m MeV}~,~
ho_0 = 0.16\,{
m fm^{-3}}~,~{
m K} = 290\,{
m MeV}$

- Realistic (complex, momentum dependent) single-particle potential ... satisfying Hugenholtz - van Hove and Luttinger theorems (!)
- Asymmetry energy: $A(k_F^0) = 34 MeV$

Fermi Liquid Theory: Quasiparticle interaction and Landau parameters J.W. Holt, N. Kaiser, W.W. Nucl. Phys. A 870 (2011) 1, Nucl. Phys. A 876 (2012) 61, Phys. Rev. C 87 (2013) 014338

C. Wellenhofer, J.W. Holt, N. Kaiser, W.W. Phys. Rev. C 89 (2014) 064009

Recent review: J.W. Holt, N. Kaiser, W.W. Prog. Part. Nucl. Phys. 73 (2013) 35

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PHASE DIAGRAM of NUCLEAR MATTER

Trajectory of **CRITICAL POINT** for **asymmetric matter**

as function of proton fraction Z/A

S. Fiorilla, N. Kaiser, W.W. Nucl. Phys. A880 (2012) 65

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... determined almost completely by **isospin** dependent (one- and two-) **pion** exchange dynamics

NEUTRON MATTER

In-medium chiral effective field theory (3-loop) with resummation of short distance contact terms (large nn scattering length, $a_s = 19 \text{ fm}$)

- agreement with sophisticated many-body calculations
 - (e.g. recent Quantum Monte Carlo computations)

PART II:

Chiral Nucleon-Meson Model and Functional Renormalization Group

S. Floerchinger, Ch. Wetterich : Nucl. Phys. A 890-891 (2012) 11

Mesonic and nucleonic particle-hole fluctuations

treated non-perturbatively using **FRG**

M. Drews, T. Hell, B. Klein, W.W. Phys. Rev. D 88 (2013) 096011

M. Drews, W.W. arXiv:1404.0882

CHEMICAL FREEZE-OUT

S. Floerchinger, Ch. Wetterich : Nucl. Phys. A 890-891 (2012) 11

Chiral nucleon - meson model in mean-field approximation

Chemical freeze-out in baryonic matter at T < 100 MeV is not associated with (chiral) phase transition or rapid crossover

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Fixing the input: some comments

• Potential
$$\mathcal{U}(\sigma, \pi) = \mathcal{U}_0(\chi) - m_\pi^2 f_\pi(\sigma - f_\pi)$$

chiral invariant part
parametrized in powers of
 $\chi = \frac{1}{2}(\sigma^2 + \pi^2)$
explicit chiral
symmetry breaking

Scalar ("sigma") field has mean-field (chiral order parameter) and fluctuating pieces. σ mass: NOT to be confused with pole in I = 0 s-wave pion-pion T matrix. Nucleon mass: $m_N^2 = 2g \chi$... in vacuum: $m_N = g f_{\pi}$

Vector fields encode short-distance NN dynamics, treated as self-consistently determined background mean fields (non-fluctuating) (not to be identified with physical ω and ρ mesons)

Effective chemical potentials $\mu_{n,p}^{\text{eff}} = \mu_{n,p} - g_{\omega} \omega_0 \pm g_{\rho} \rho_0^3$ Relevant quantities: $G_{\rho} = \frac{g_{\rho}^2}{m_V^2}$, $G_{\omega} = \frac{g_{\omega}^2}{m_V^2} \iff \text{contact terms in ChEFT}$

Parameters: 2 coefficients in U_0 , $m_\sigma \simeq 0.8 \, GeV$, $G_\rho \sim G_\omega/4 \simeq 1 \, fm^2$ determined by tuning nuclear matter properties and symmetry energy

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Chiral nucleon - meson model beyond mean-field - Renormalization Group strategies -

M. Drews, T. Hell, B. Klein, W.W. Phys. Rev. D 88 (2013) 096011

• Fluctuations: Wetterich's RG flow equations • **full propagator** $k \frac{\partial \Gamma_k}{\partial k} = \sum = \frac{1}{2} \operatorname{Tr} \frac{k \frac{\partial R_k}{\partial k}}{\Gamma_k^{(2)} + R_k}$ • **regulator**: $R_k(p^2) = (k^2 - p^2) \theta(k^2 - p^2)$ • **C**. Wetterich: Phys. Lett. B 301 (1993) 90 D.F. Litim, J.M. Pawlowski: JHEP 0611 (2006) 026 J.-P. Blaizot, A. lpp, R. Mendez-Galain, N.Wschebor: NP A784 (2007) 376

Thermodynamics nucleons pions

$$k \partial_k \bar{\Gamma}_k(T,\mu) = \left(\bigotimes + \bigotimes \right) \Big|_{T,\mu_p,\mu_n}$$

 $-\left(\bigotimes + \bigotimes \right) \Big|_{T=0,\mu=\mu_0 (=m_N - E_0/A)}$

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Γ_{kas} is national for the second se

Flow equations in practice

 k_m son magalanti The Yukawa (4.900)

Results : Liquid - Gas Transition

- symmetric nuclear matter -

M. Drews, T. Hell, B. Klein, W.W. Phys. Rev. D 88 (2013) 096011

note: close correspondence between (perturbative) in-medium ChEFT and (non-perturbative) FRG results

 Single additional parameter:
 coupling strength of isovector-vector field / contact term fixed by symmetry energy E(sym) = 32 MeV

Chiral Order Parameters

Comparison of chiral effective field theory and model FRG results

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In-medium pion mass

Contact with phenomenology :

compare with s-wave pion-nuclear optical potential from pionic atoms

 Good agreement of FRG calculation with empirical in-medium pion mass shift, both in sign and magnitude

PART III: Short digression on **Neutron Stars** outer crust 0.3-0.5 km ions, electrons inner crust 1-2 km 0.5-2.0 ρ₀ electrons, neutrons, nuclei >2.0 p outer core ~ 9 km neutron-proton Fermi liquid few % electron Fermi gas inner core 0-3 km quark gluon plasma?

NEUTRON STARS and the **EQUATION OF STATE** of **DENSE BARYONIC MATTER**

J. Lattimer, M. Prakash: Astrophys. J. 550 (2001) 426 Phys. Reports 442 (2007) 109

Mass-Radius Relation

New constraints from NEUTRON STARS

PSR J1614+2230

 $\mathbf{M} = 1.97 \pm 0.04~M_{\odot}$

ECT*

J.Antoniadis et al. Science 340 (2013) 6131

PSR J0348+0432

 $\mathbf{M} = \mathbf{2.01} \pm \mathbf{0.04} \,\, \mathrm{M_{\odot}}$

CONCLUSIONS

FRG provides non-perturbative approach to Nuclear Chiral Thermodynamics

from symmetric to asymmetric nuclear matter and neutron matter

fluctuations beyond mean field include important two-pion exchange mechanisms and low-energy nucleonic particle-hole excitations

🖿 l st order phase transition: Fermi liquid \leftrightarrow interacting Fermi gas

No indication of **first-order chiral phase transition**

- fluctuations work against early restoration of chiral symmetry
- New constraints from neutron stars for the equation-of-state of dense & cold baryonic matter:
 - Mass radius relation: **stiff equation of state** required ! **No** ultrahigh densities $(arrho_{\max} \sim 5 \, arrho_0)$

Conventional (nucleon-meson, "non-exotic") EoS fulfills constraints